

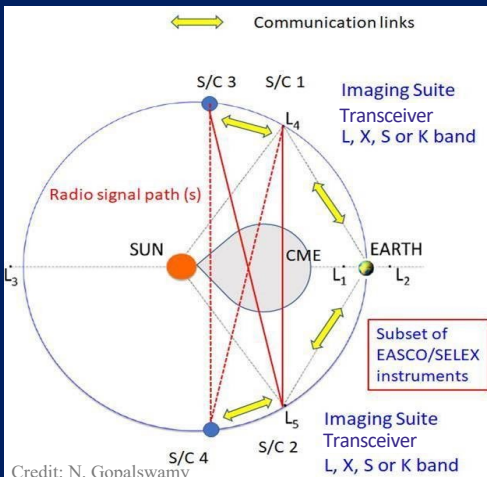
Faraday Effect Tracker of Coronal and Heliospheric Structures (FETCH): Instrument Concept



M. Kenny (CU Boulder, SwRI) and the FETCH Team*



Multiview Observatory for Solar Terrestrial Science (MOST): Mission Concept



MOST Instruments
Magnetic and Doppler Imagers (MaDI)
Inner Coronal Imager in EUV (ICIE)
Hard X-ray Imager (HXI)
White-light Coronagraph (WCOR)
Faraday Effect Tracker of Coronal and Heliospheric structures (FETCH)
Heliospheric Imager with Polarization (HIP)
Radio and Plasma Wave Instrument (M/Waves)
Energetic Particle Detector (EPD)
Solar Wind Magnetometer (MAG)
Solar Wind Plasma Instrument (SWPI)

- **4 spacecraft (s/c):**
 - 2 large s/c: at Lagrange points L4 & L5
 - 2 smaller s/c: one ahead of L4, one behind L5
- **Radio suites on all 4 s/c**

Motivations for MOST and FETCH

- Solar disturbances and wind are magnetic by nature
- Observations of **magnetic field structure** in solar wind, coronal mass ejections (CMEs), stream interaction regions (SIRs), and shocks **upstream of Earth and L1**
- Multiple spacecraft → multiple viewpoints / lines-of-sight
- **Absence of ionosphere** in space-based FR

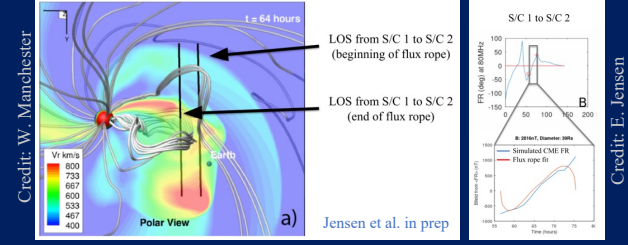
FETCH CME Science

- **Will probe magnetic field structure and CME evolution using Faraday rotation (FR) from 15R_☉ to 0.5 AU**
 - Four lines-of-sight upstream of Earth, entirely outside the ionosphere
- Radio frequency bandpass: **1-100 MHz**

Science Objectives	Required Measurements	
	Physical Parameters	Observables
Within CME structure, determine the distribution in the plane-of-sky bulk velocity, average density, & magnetic field at varying distances from the Sun	CME , from Faraday rotation (FR), dFR, total electron content (TEC), and dTEC	Stokes parameters
Analyze force balance within CMEs	CME <N>, from TEC and dTEC	Time delay (group velocity) & frequency fluctuations

Significance of FETCH Measurements:

- **Space weather implications:** magnetic topology of solar wind and CMEs upstream of Earth
- **Present modeling efforts:**
 - 2005 CME event (left figure), detectable fluctuations (e.g. Alfvén waves), multi-LOS FR (right figure), transmitter-receiver antenna designs



Broader Impacts

- **Collaboration and interdisciplinary efforts:**
 - Complementary MOST instruments: HIP, WCOR
 - Context measurements: Global Oscillation Network Group (GONG) and ngGONG
- **Reach out; get involved:**
 - FETCH: Dr. Jensen (ejensen@psi.edu)
 - MOST: Dr. Gopalswamy (natchimuthuk.gopalswamy-1@nasa.gov)

Acknowledgments

This material is based upon work supported by the National Science Foundation Graduate Research Fellowship under Grant DGE 1650115. Any opinion, findings, and conclusions or recommendations expressed in this material are those of the authors(s) and do not necessarily reflect the views of the National Science Foundation.

* N. Gopalswamy (GSFC, Mission Lead), E.A. Jensen (PSI, Instrument Lead), S.D. Bale (UC-B), T. Bastian (NRAO), S. Fung (GSFC), L. Jian (GSFC), L. Li (GSFC), T. Nieves-Chinchilla (GSFC), L.B. Wilson (GSFC), J. Lazio (JPL), W. Manchester (UMich), D. Wexler (UML), A. Pevtsov (NSO)