

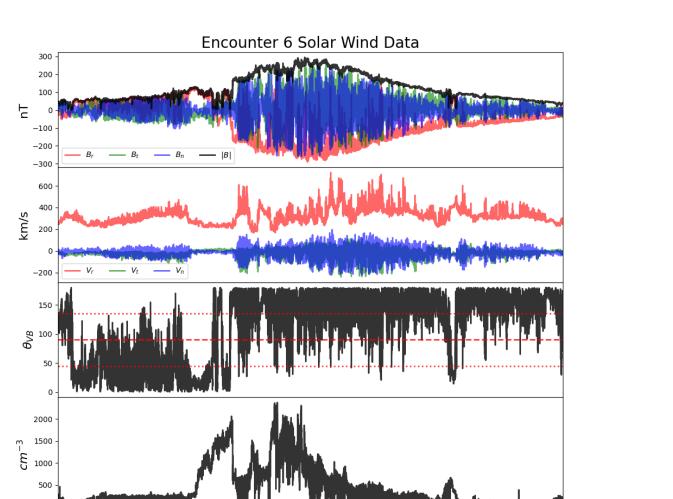


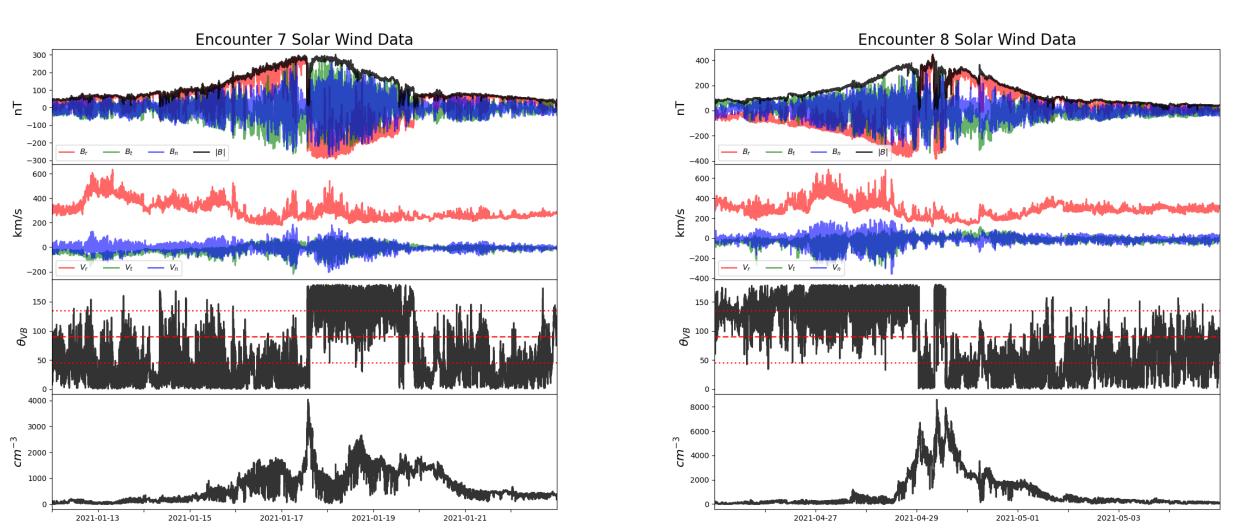
1. OVERVIEW

- Alfvénic turbulence is developed by the interaction of inward/outward propogating Alfvén waves.
- *In situ* measurements show predominant preferential outward direction.
- Alfvén cascade cannot devlop without the presence of inward propagating wages.
- Theoretically, $\mathbf{z}^{\pm} = \mathbf{u} \pm \mathbf{v}_A$ where \mathbf{u} and $\mathbf{v}_A = \delta \mathbf{B}/4\pi \rho$ are fluctiations, represent outward/inward Alfvén waves.
- In highly outward Alfvénic flows, z⁻ spectrum at low frequency is convected structures, while high end trails off to white noise. (Wang et. al. 2018)
- PSP will probe across the Alfvén critical point.

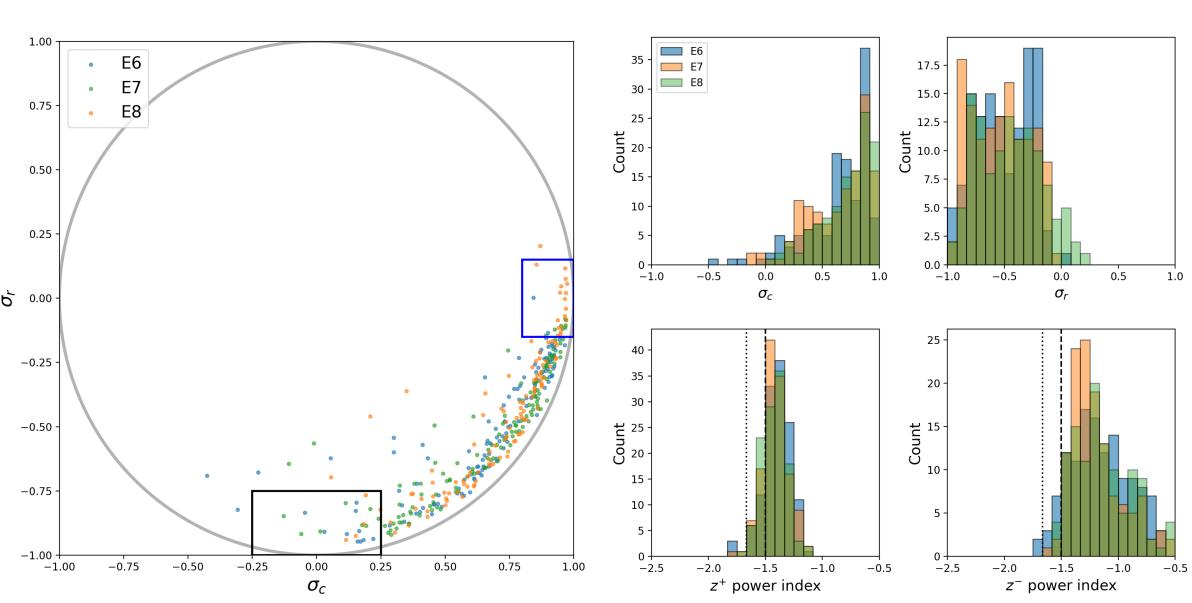
4. CHARACTERIZING PSP ENCOUNTERS 6, 7, AND 8

• So what do the bulk properties of PSP Encounters 6, 7, and 8 look like?





• We can already see that each period is predominately outward Alfvén waves. To see this better we can look at the residual energy and cross-helicity.



A Measure of Alfvénic Turbulence during PSP encounters 6, 7, and 8

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2. MOTIVATION

- Can we extract out the true power of Alfvén waves?
- What information can we gather at various cross-helicities?

3. TURBULENCE TOOLS

• Rectified Elsässer variables:

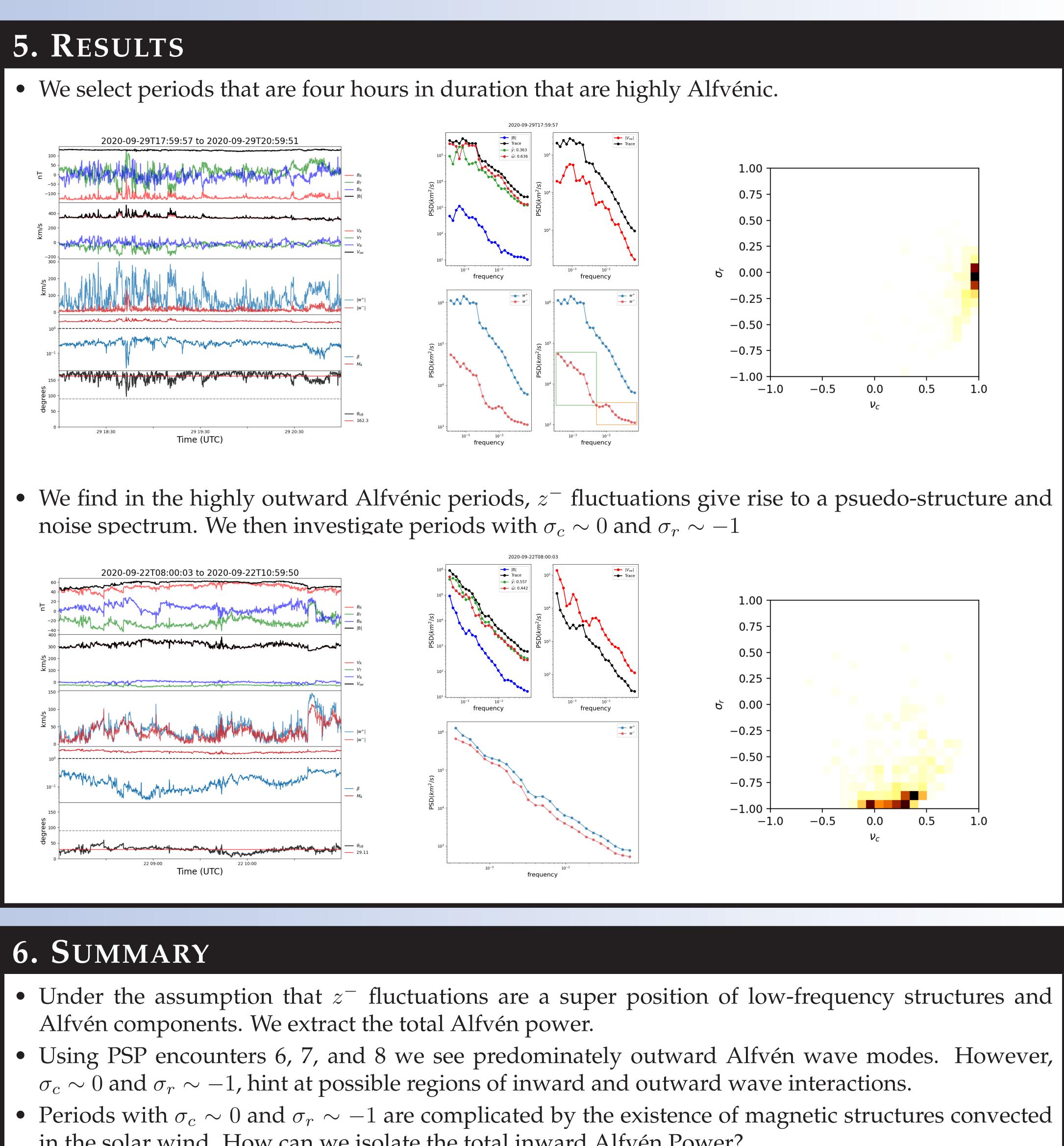
$$\mathbf{w}^{\pm} = \begin{cases} \mathbf{z}^{\pm} & \text{if } \operatorname{sgn}(B_{0r}) = -1 \\ \mathbf{z}^{\mp} & \text{if } \operatorname{sgn}(B_{0r}) = 1 \end{cases}$$
(1)

• Residual Energy:

$$\sigma_r = \frac{2\langle \mathbf{z}^+ \cdot \mathbf{z}^- \rangle}{\langle z^{+2} \rangle + \langle z^{-2} \rangle} = \frac{\langle u^2 \rangle - \langle b^2 \rangle}{\langle u^2 \rangle + \langle b^2 \rangle}$$

• Cross Helicity:

$$\sigma_{c} = \frac{\langle z^{+2} \rangle - \langle z^{-2} \rangle}{\langle z^{+2} \rangle + \langle z^{-2} \rangle} = \frac{2 \langle \mathbf{u} \cdot \mathbf{b} \rangle}{\langle u^{2} \rangle + \langle b^{2} \rangle}$$





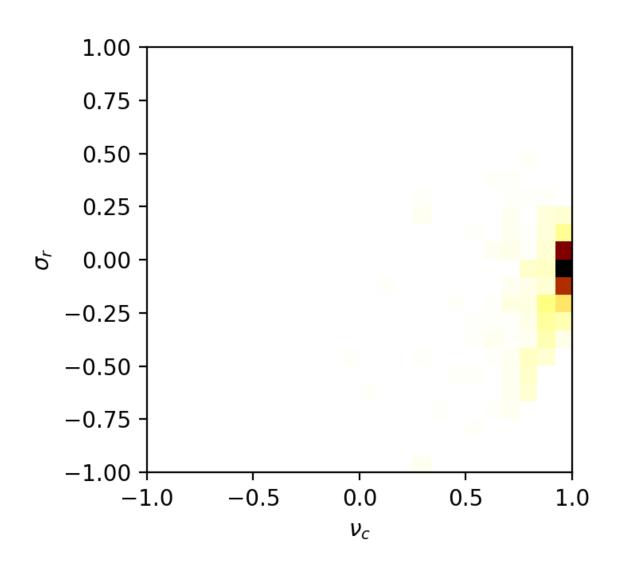
in the solar wind. How can we isolate the total inward Alfvén Power? • We see total power is perpendicular to the background magnetic field for both the highly Alfvénic and $\sigma_c \sim 0$ and $\sigma_r \sim -1$.

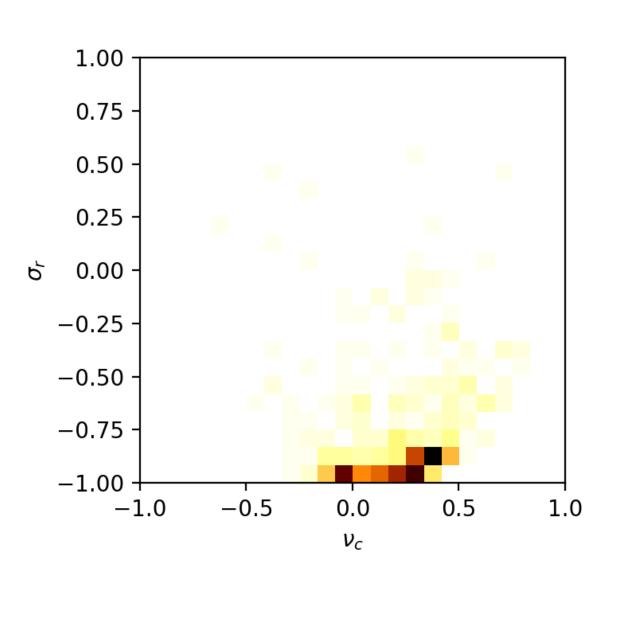
7. ACKNOWLEDGMENT

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8. CONTACT INFORMATION

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