

# Dynamics of Cosmic Ray Muon Flux under the Influence of Earth's Atmosphere, Solar Activity, and Geomagnetic Turbulence

A. Mubashir\*, A. Ashok\*\*, A.G. Bourgeois\*\*, Y.T. Chien\*, M. Connors\*, E. Potdevin\*, X. He\*, P. Martens\*, A. Mikler\*\*, A.G.U. Perera\*, V. Sadykov\*, M. Sarsour\*, D. Sharma\*, C. Tiwari\*\*,†

\* Department of Physics and Astronomy, \*\* Department of Computer Science, † Department of Geosciences, Georgia State University

## Introduction

- The solar activity, the state of the interplanetary space, and the Earth's magnetosphere and atmosphere are collectively responsible for the intensity of secondary cosmic ray particles being detected by ground-based cosmic ray detectors, as highlighted in (Fig. 1)

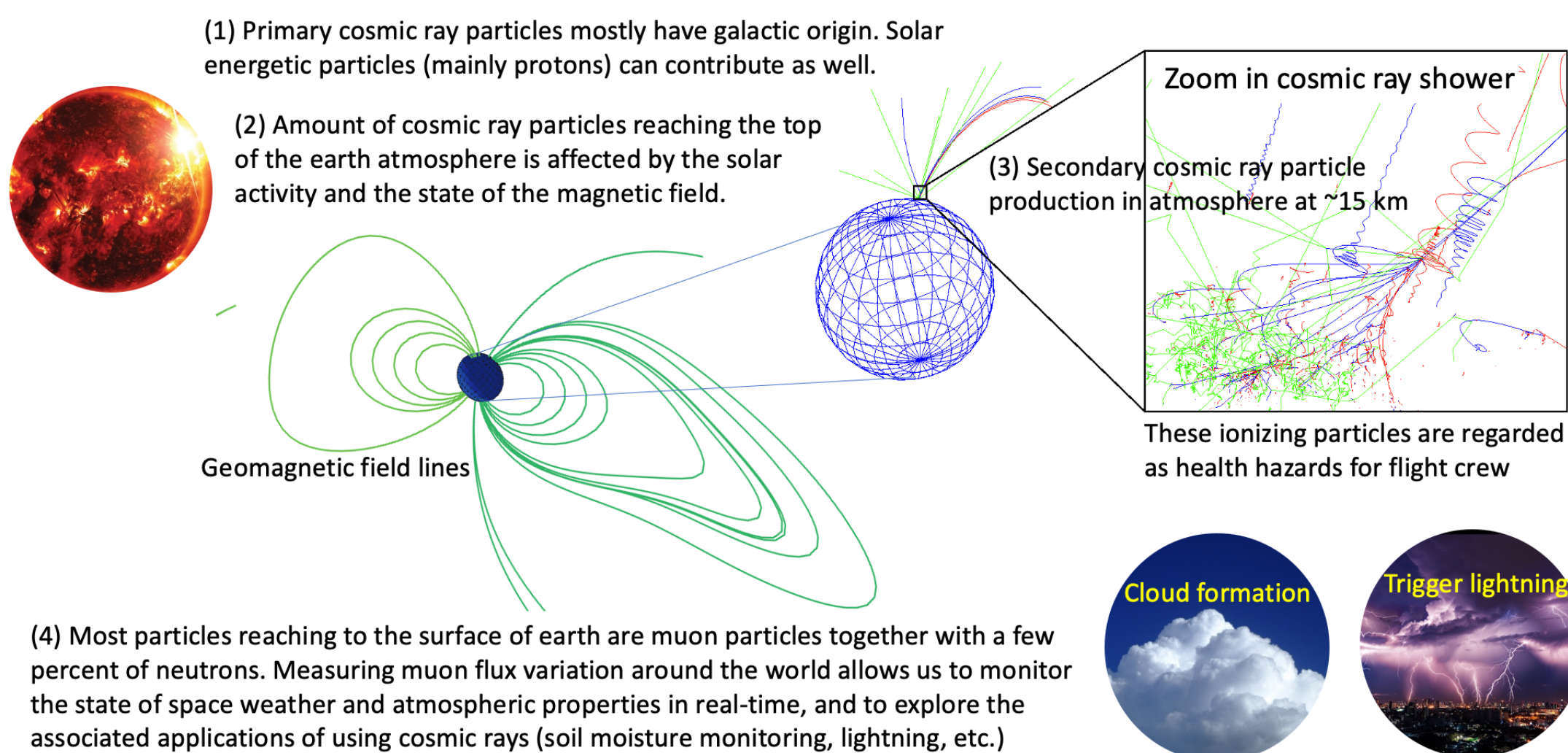


Figure 1: Cosmic rays: from space to Earth

- Key challenge: building efficient and affordable network of cosmic ray muon detectors capable of providing the sensitivity accurately to the variations of both space and terrestrial weather patterns.

## Motivation

- A network of cosmic ray muon detectors that can provide real-time monitoring of Earth's atmospheric parameters, solar activities and geomagnetic disturbances.
- Possibilities for widespread deployment, facilitating extensive data collection and analysis in various geographical locations.

## Detector setup

- Desktop cosmic ray muon detector design:
  - Cosmic ray muon detector dimension and configuration



- Detector acceptance study using GEANT4 simulation toolkit (Fig. 2).

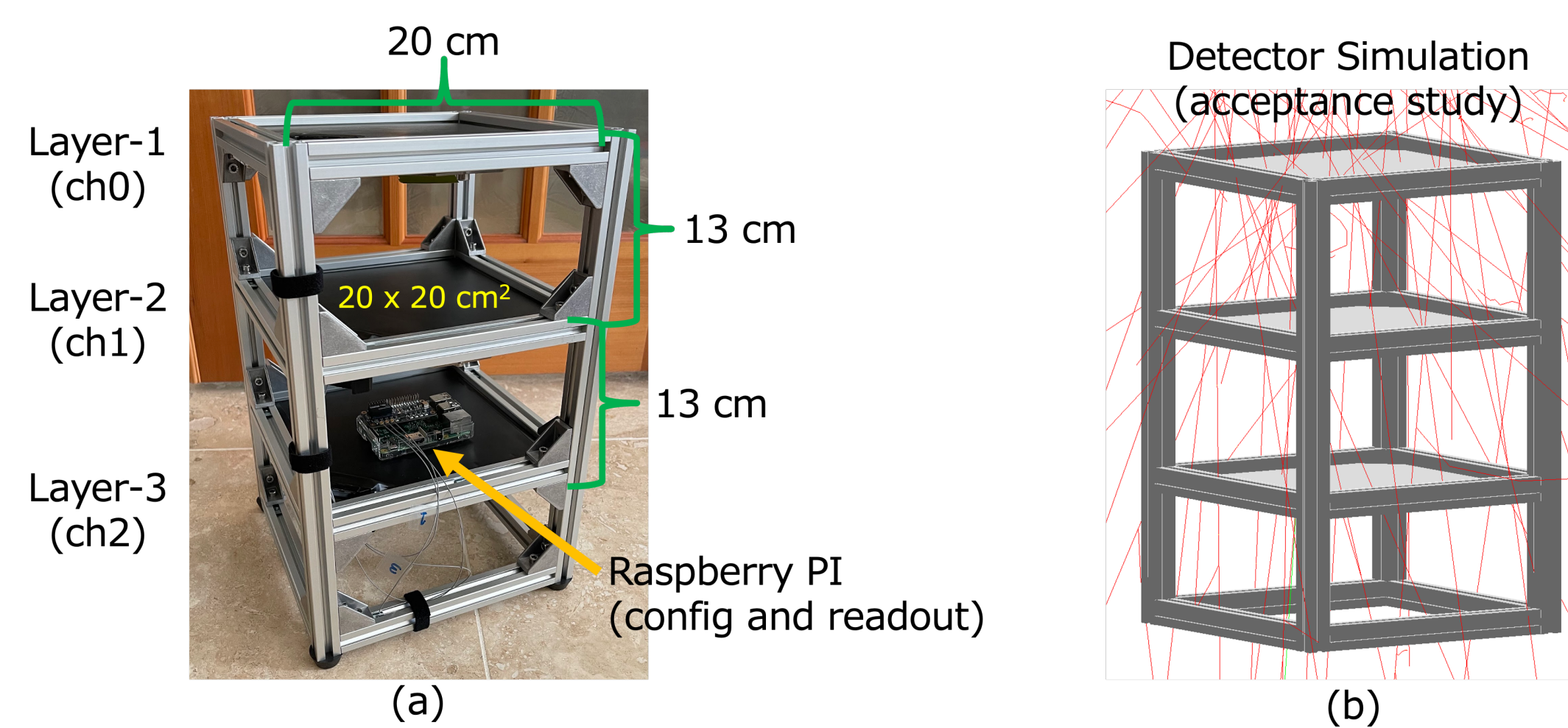


Figure 2: Muon detector used in this study

## Atmospheric effect on muon flux

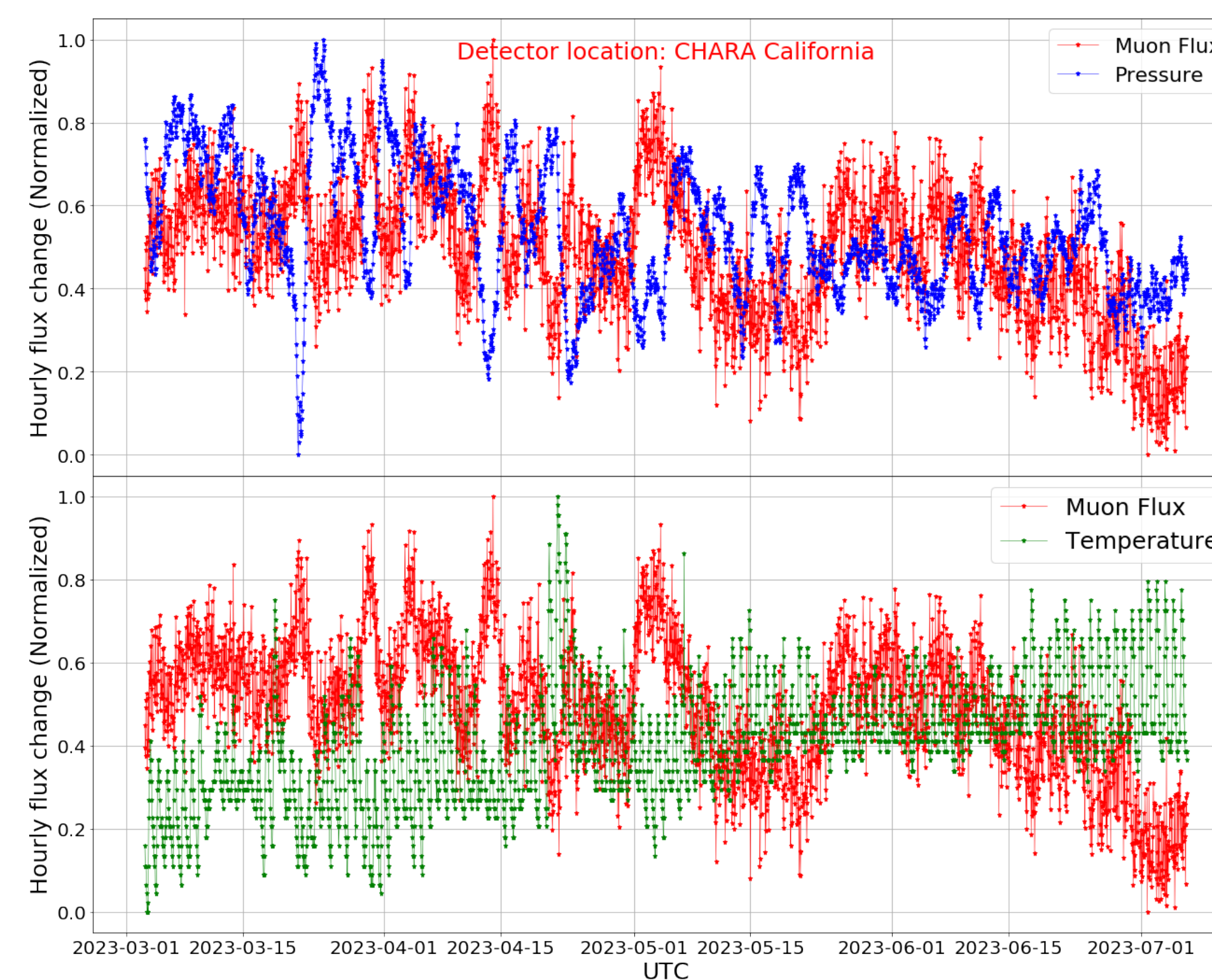


Figure 3: Muon flux vs pressure and temperature

- Ensuring the accuracy of muon data by applying temperature and pressure corrections.

## Cross comparison of cosmic ray data and space weather Activity

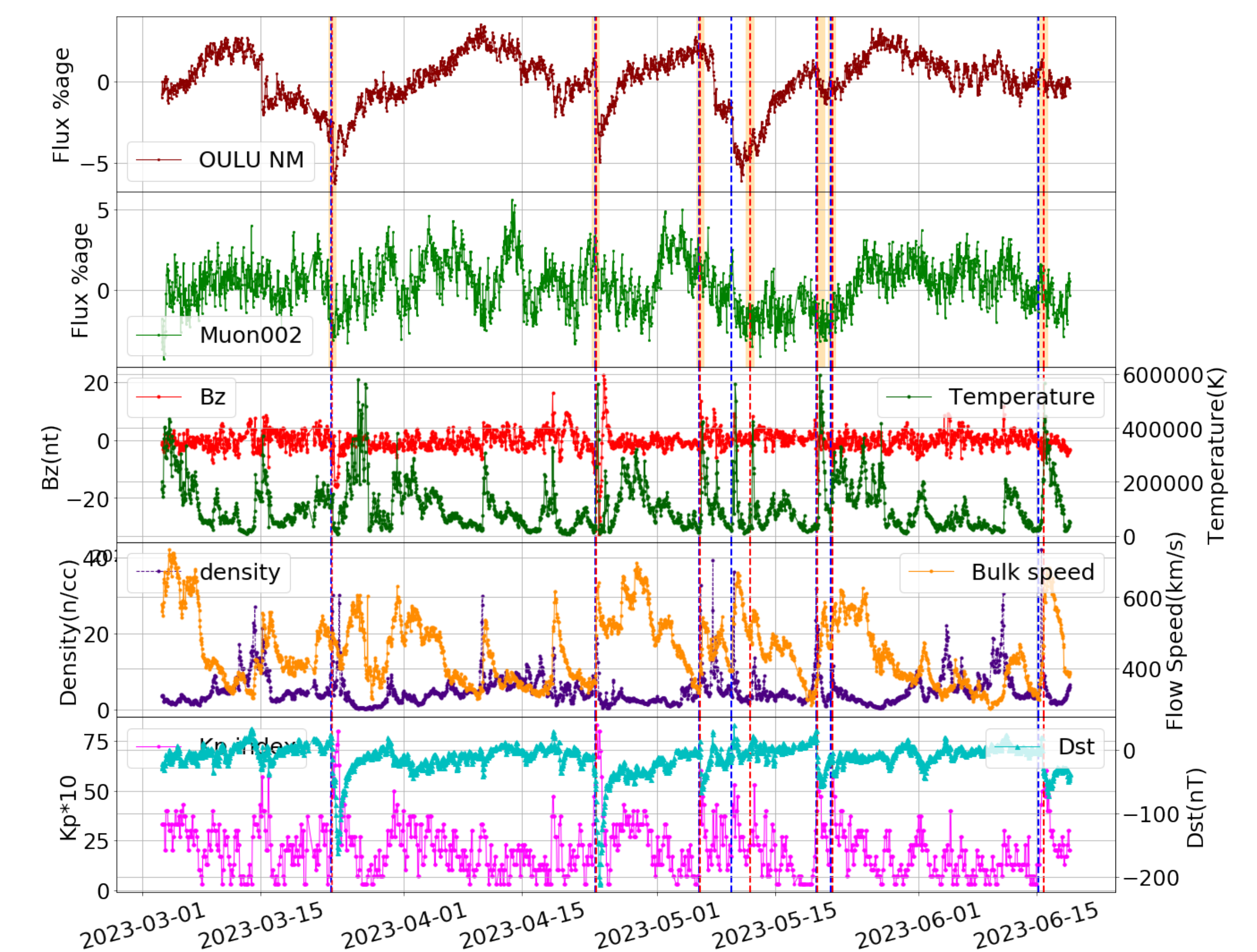


Figure 4: Cosmic ray flux and space parameters

## Network expansion

- Installed two detectors in Sri Lanka in March 2023 Building more detectors to be distributed in other countries.

## Summary

- Oulu and CHARA detectors show similar response to space weather events.
- Pearson coefficients between flux and space parameters are weakly but statistically significant ( low p-values i.e  $10^{-34}$ ,  $10^{-25}$ ).
- Decreasing trend in cosmic ray flux before the geomagnetic storm.

## Acknowledgement

This work is supported by Georgia State University's RISE program.

