

Nature and Properties of Electrostatic Solitary Waves in the Earth's Magnetosheath

Zubair Shaikh*¹ and Ivan Vasko²

(Email: zshaikh@berkeley.edu)

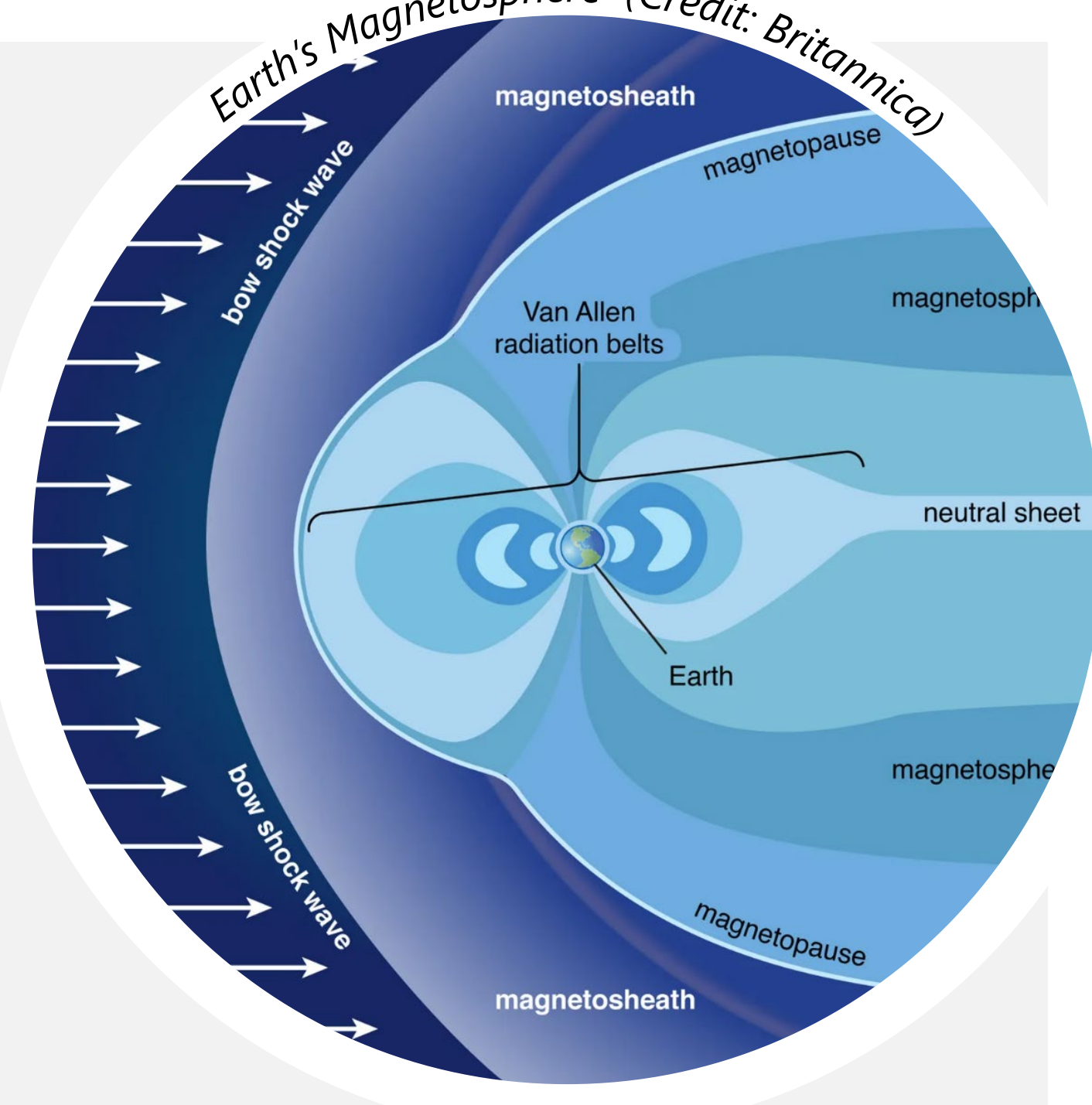
1. Space Sciences Laboratory, UC Berkeley, CA, USA; 2. William B. Hanson Center for Space Sciences, UTD, Richardson, TX, USA



Website

Background and Motivation

- **Electrostatic solitary waves (ESWs)** are **nonlinear**, **localized**, **bipolar** electric field structures propagating **parallel/anti-parallel** to the background magnetic field¹⁻¹⁵.
- ESWs are categorised into:
 - Solitons: Electron/Ion solitons,
 - **Phase-space holes**: Electron/Ion holes
- ESWs plays an important role in plasma transport, particle scattering, and energy exchange processes^{1,2,15}.



ESWs in Earth's Magnetosheath

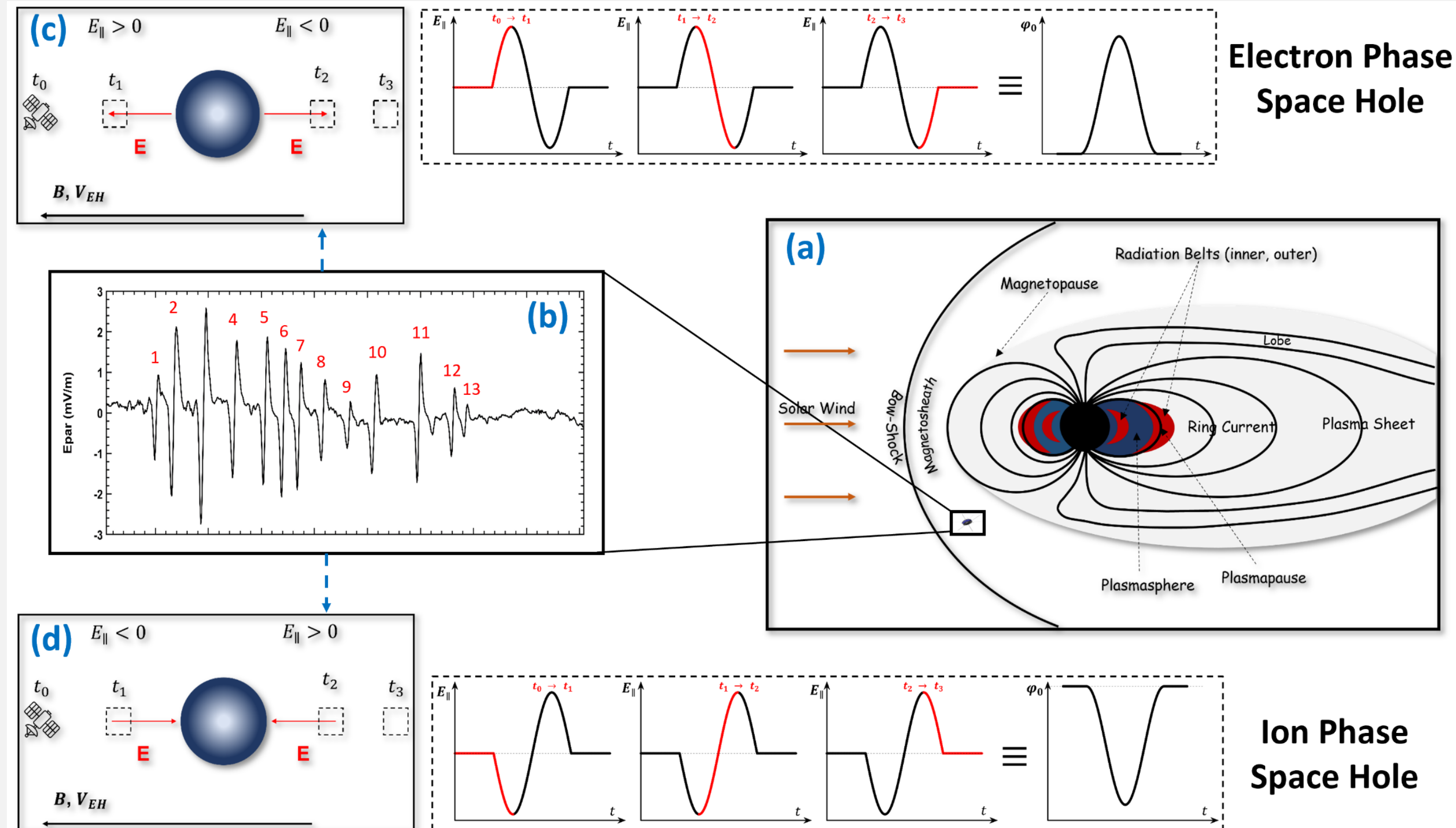


Figure 1: Schematics of (a) Solar wind - Earth's magnetosphere interaction; (b) electrostatic solitary waves within the Earth's Magnetosheath; (c) Electron hole; and (d) Ion hole.

- ESWs were observed Magnetosheath...*Kojima+1997; Pickett+(2003, 2005)*. } **Characteristic were unknown**
 - *Graham+2016* observed ESWs with positive polarity (**electron holes**).
 - *Holmes+2017* observed ESWs with positive polarity (**ion holes**).
 - *Shaikh+2024* observed ESWs with positive polarity (**slow electron holes**).
- Inconsistent between these studies motivated the present study**

Science objectives

- Which Solitary waves are often present within the Magnetosheath?
- What are the properties and generation mechanism of solitary waves?
- How does turbulence level affects characteristics and occurrence of ESWs?

Data and Methodology

- **MMS S/C:** FPI (Ions:150ms, electrons:30ms), FGM (128Hz), EDP (8192 Hz), and SCM (8192 Hz).
- **Methodology**^{13,15-16}:
 - Minimum Variance Analysis (MVA)
 - Voltage Interferometry Analysis
- **Assuming locally planar and one dimensional solitary wave**¹³⁻¹⁸
 - Solitary wave speed; $V_s = \frac{\partial E_{ij}}{\partial t_{ij}}$ here, \hat{E}_{ij} is polarization direction
 - Electrostatic potential; $\varphi = \int E_L V_s dt$
 - Spatial scale of solitary structure; $x = \int V_s dt$; with $x = 0$, corresponding to $E_L = 0$.

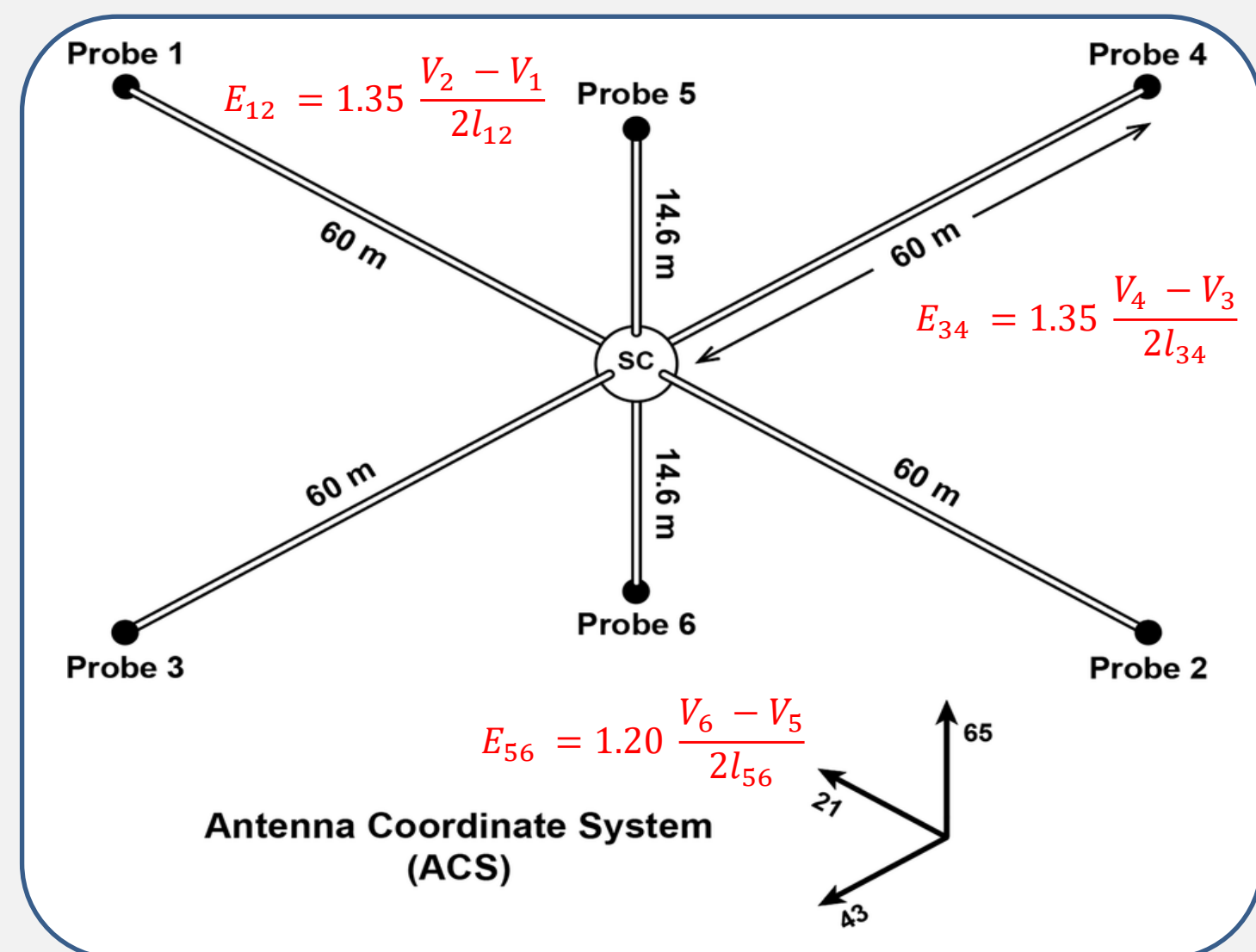


Figure 2: The schematics of the double probe instrument aboard on each MMS spacecraft²

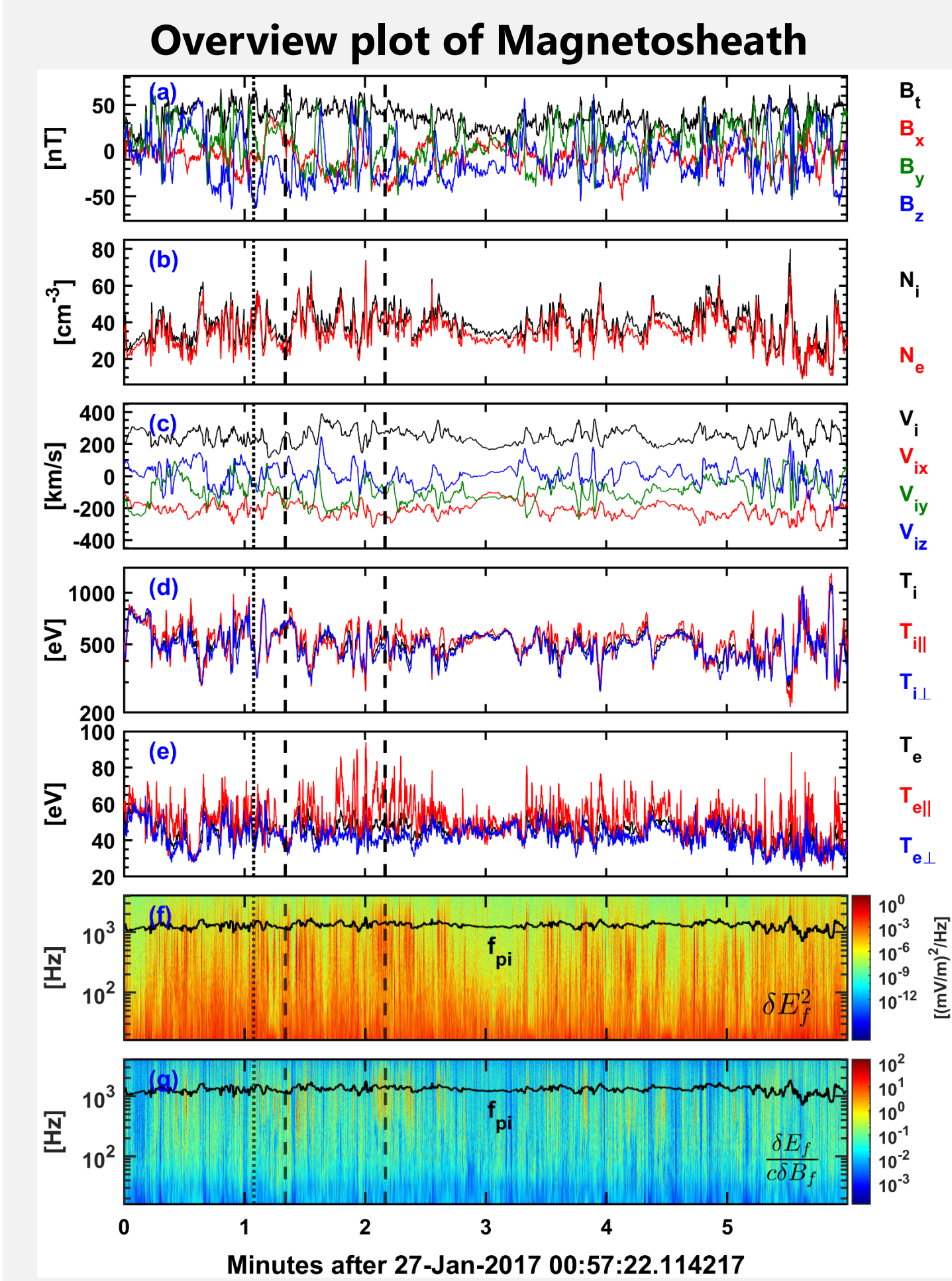


Figure 3: Magnetosheath overview on January 28, 2017.

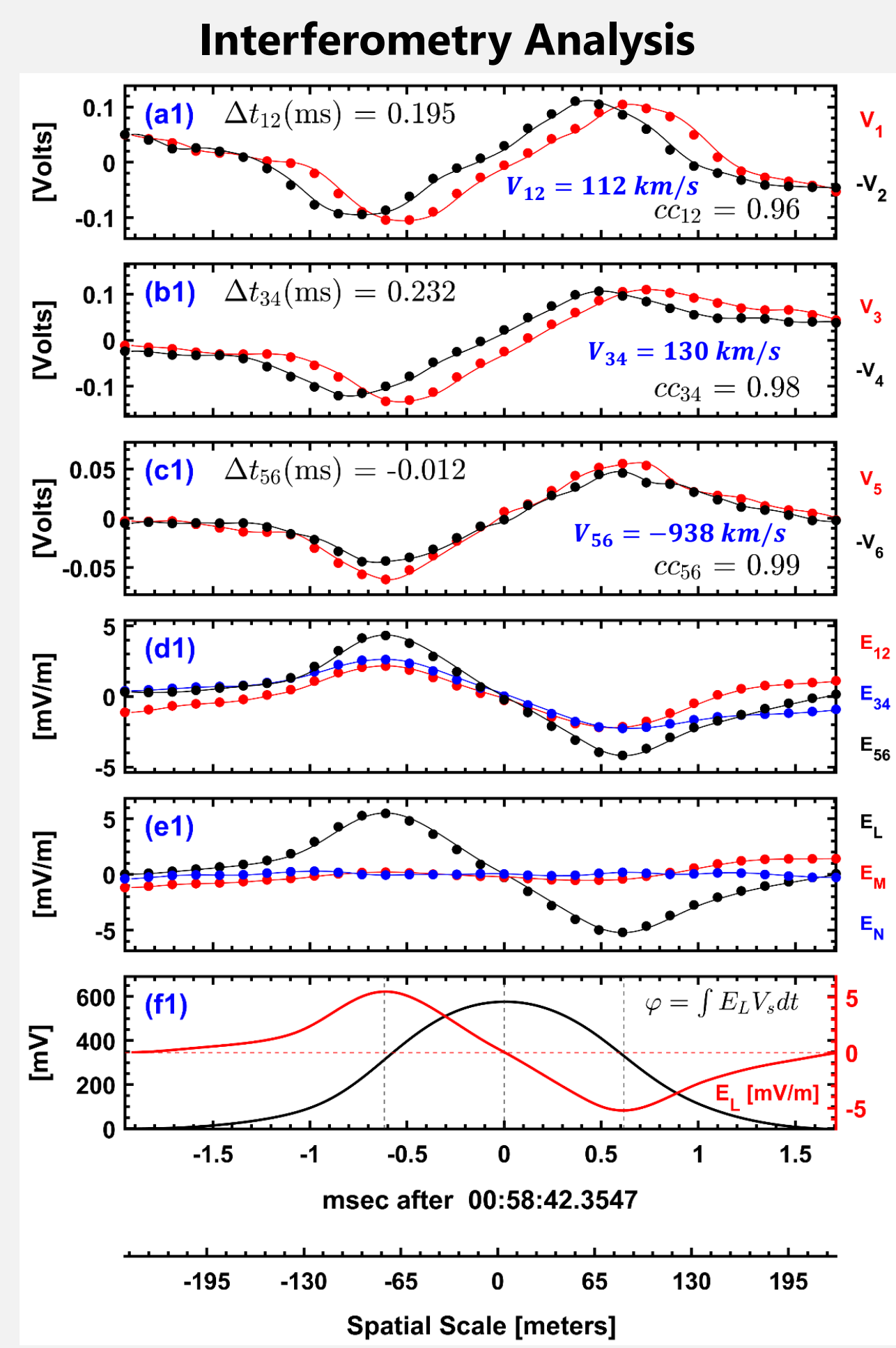


Figure 4: An example of ESW with positive potential.

Selection Criteria

1. At least **one pair of voltage probe** must show:
 - $cc \geq 0.85$
 - $\delta t \geq 0.06 \text{ msec}$
2. $E_L \gg E_M, E_N$ which is equivalent to $\frac{\lambda_{max}}{\sqrt{\lambda_{int}^2 + \lambda_{min}^2}} > 5$

Table 1. List of Magnetosheath, duration, turbulent level, number of observed IHS/EHS by MMS1 & MMS2.

Magnetosheath Interval	Duration ΔT (min)	Turbulence $\delta B_{rms}/B_0$	#IHS		#EHS	
			MMS1	MMS2	MMS1	MMS2
Oct 31, 2015; 10:22:27 to 10:27:00	04.55	0.64	134	194	54	91
Jan 27, 2017; 00:55:03 to 00:56:58	01.92	1.05	105	122	277	269
Jan 27, 2017; 00:57:23 to 01:03:13	05.83	2.46	191	362	103	241
Total			430	678	434	601

IHS = 1108 EHS = 1035

Analysis and Results

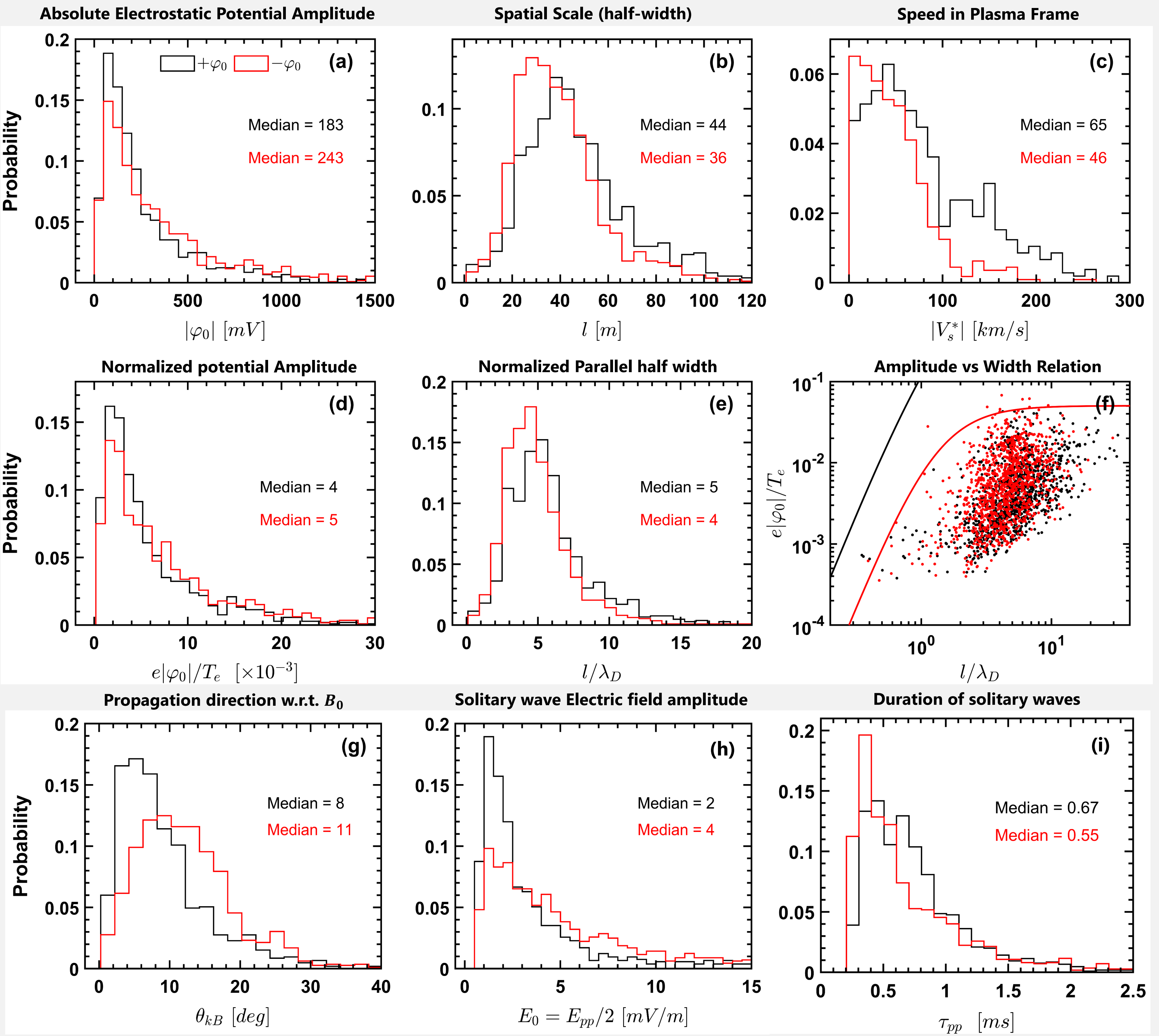


Figure 5: Distributions of solitary waves properties (**black: positive potential structures, red: negative potential structures**).

Ion velocity distribution functions (IVDFs)

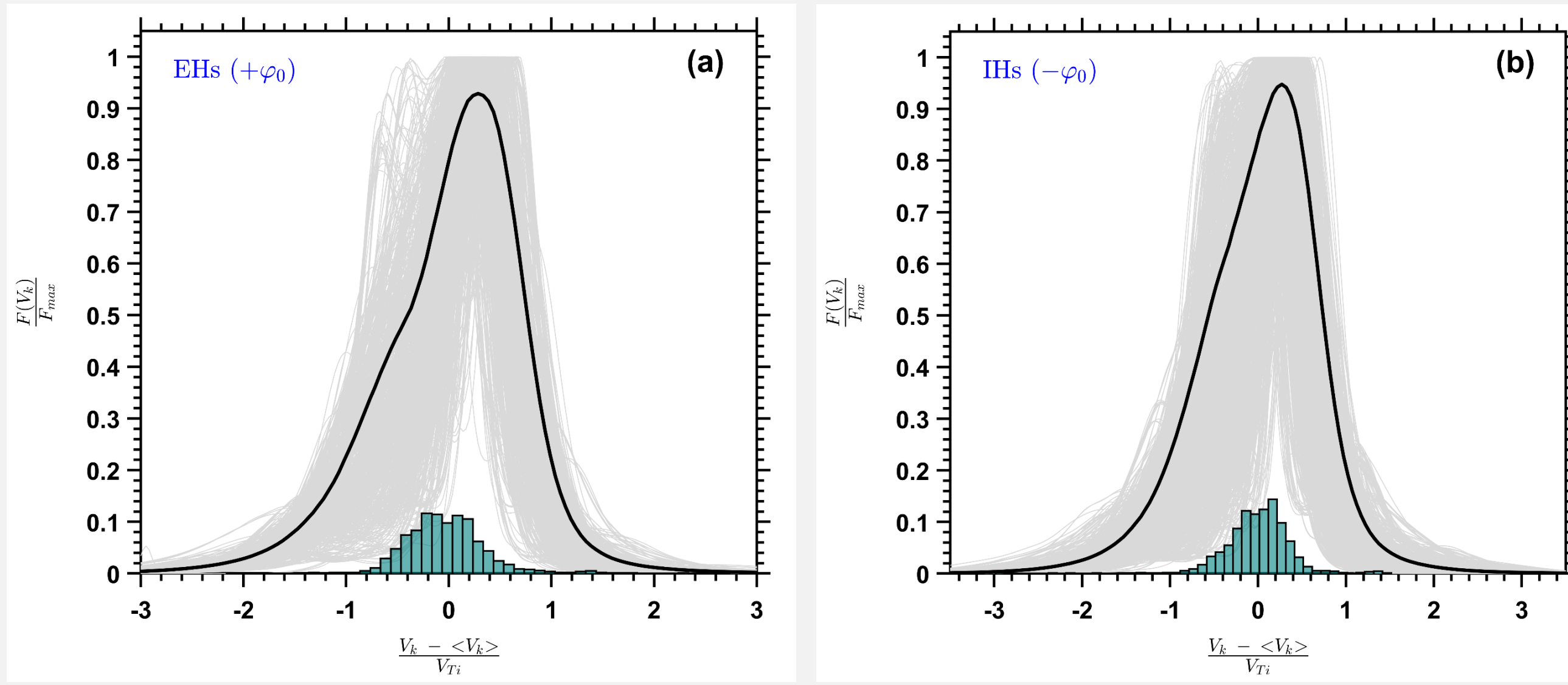


Figure 6: 1D reduced IVDFs $F(V_k)$ in the plasma frame collected over 150 ms around the solitary waves. The grey color distribution is for individual solitary waves, whereas the overlapped black curves present average distributions. Each $F(V_k)$ is normalized with peak value F_{max} and translated into the plasma frame and normalized to ion thermal speed V_{Ti} . The histograms present normalized probability distributions of solitary wave speed in the plasma frame.

Summary and Conclusions

Statistically analysis >2,100 ESWs within the three Earth's Magnetosheath crossings having different turbulence amplitude:

- About **48%** of the ESWs are **electron holes (EHs)**, whereas **52%** are **ion holes (IHS)**.
- **EHs and IHS have similar properties** such as; spatial scales of a few Debye lengths, plasma frame velocities within about 200 km/s, and potential amplitudes < 2 V.
- **EHs and IHS are slow**, their plasma frame speeds are much smaller than local electron thermal speed, but comparable with local ion thermal speed.
- **The cluster of EHs are statistically well-separated from clusters of IHS**, strongly indicates that the solitary waves of different polarities are highly likely produced by instabilities operating separately in time or space.
- Characteristics & occurrence of EHs/IHS are **insensitive to local turbulence intensity**.

Time intervals between EHs/IHS

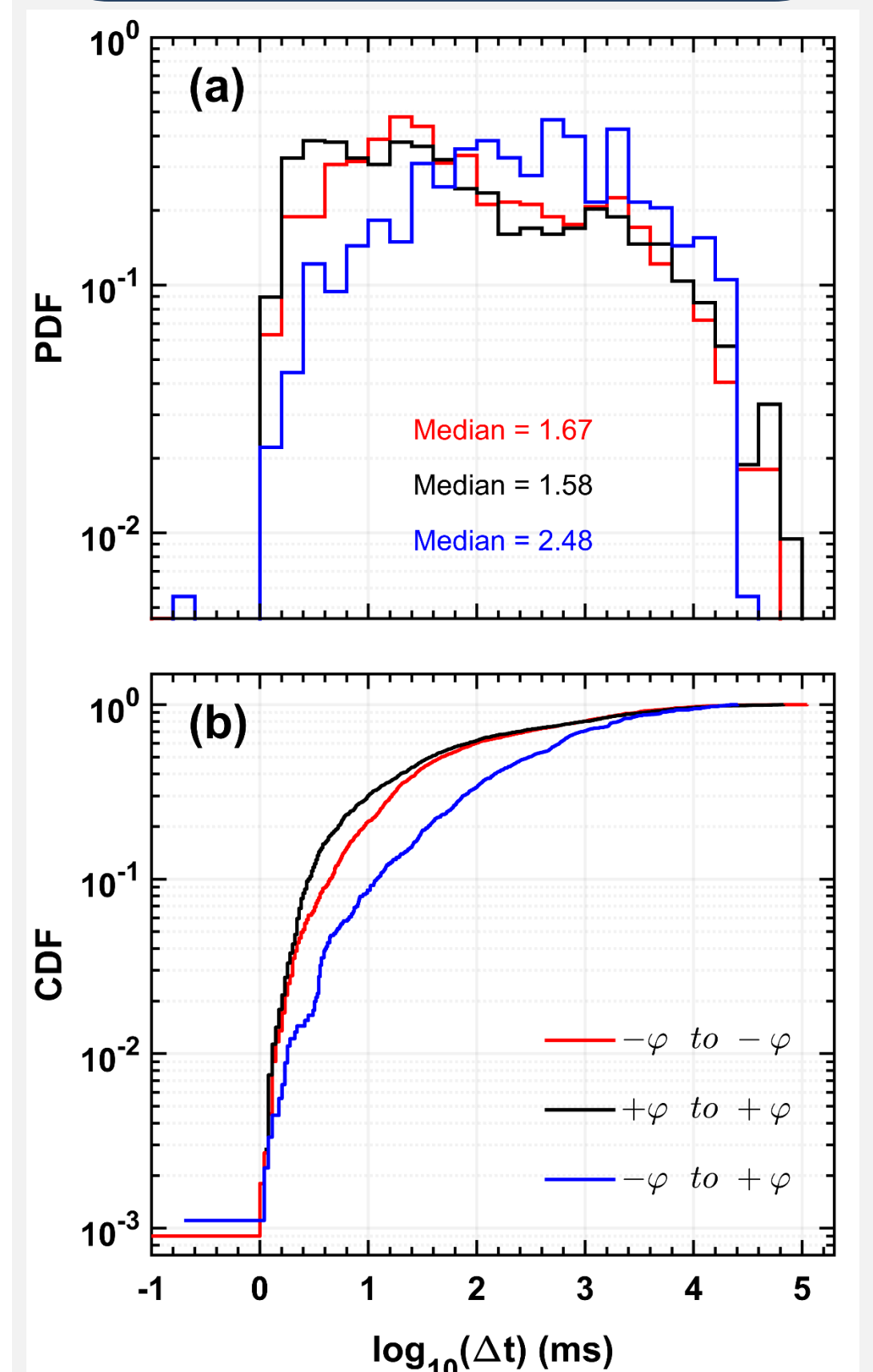


Figure 7: The probability and Cumulative distribution functions of time intervals between sequentially observed EHs and IHS.

References

1. Schamel, H. 1986, Physics reports, 140(3), 161-191.
2. Graham et al., 2016, JGR, 121(4), 3069-3092.
3. Hutchinson, I. H. 2017, POP, 24(5), 055601.
4. Mozer et al., 2018, PRL, 121(13), 135102.
5. Vasko et al., 2018, GRL, 45(12), 5809-5817.
6. Vasko et al., 2018, POP, 25(7), 072903.
7. Holmes et al., 2018, JGR, 123(1), 132-145.
8. Wang et al., 2020, ApJL, 889(1), L9.
9. Vasko et al., 2020, Frontiers in Physics, 8, 156.
10. Kamaladinov et al., 2021, PRL, 127(16), 165101.
11. Wang et al., 2021, JGR, 126(7), e29357.
12. Kamaladinov et al., 2022, POP, 29(9).
13. Steinwall et al., 2022, JGR, 127(3), e2021JA030143.
14. Wang et al., 2022, GRL, 49(8), e2022GL079191.
15. Shaikh et al., 2024, JGR, 129, e2023JA032059.

Acknowledgement

The work of Z.S. and I.V. was supported by NASA Grant 80NSSC20K1325 and NSF Grant 2026680. The work of I.V. was also supported by NASA Grant 80NSSC22K1634. We thank MMS team for providing data.

Name: Dr. Zubair Shaikh
Email: zubairshaikh584@gmail.com
Lab: SSL, UC Berkeley, CA, USA